

281(1) : Rayleigh Jeans Density of States and Reduction to One Photon Theory.

The Rayleigh Jeans density of state is based on the wave equation:

$$\nabla^2 \phi = 0 \quad - (1)$$

and does not consider photon mass and the $\underline{b}^{(3)}$ field.

Eq. (1) is an ECE wave equation without photon mass.

The Rayleigh Jeans density of states is

$$\phi = \exp(i(\omega t - \underline{k} \cdot \underline{r})) \quad - (2)$$

$$\underline{k} = k_x \underline{i} + k_y \underline{j} + k_z \underline{k} \quad - (3)$$

where:

$$\omega^2 = c^2 (k_x^2 + k_y^2 + k_z^2) \quad - (4)$$

with:

$$k_x = n_1 \frac{\pi}{L}, \quad k_y = n_2 \frac{\pi}{L}, \quad k_z = n_3 \frac{\pi}{L} \quad - (5)$$

where

$$V = L^3 \quad - (6)$$

is the volume of radiation in a cube of side L. As pointed out by Jeans this is the amount of a sphere. Therefore is the Rayleigh Jeans density of states:

$$\omega^2 = \frac{c^2 \pi^2}{L^3} (n_1^2 + n_2^2 + n_3^2) \quad - (7)$$

where:

$$n^2 = n_1^2 + n_2^2 + n_3^2 \quad - (8)$$

The number of photons in the octant of the sphere is:

$$N = \frac{1}{8} \left(\frac{4}{3} \pi n^3 \right) \quad - (9)$$

$$= \frac{1}{6\pi^2} \frac{\omega^3}{c^3} L^3$$

So the number of photons in the volume V is:

$$\frac{N}{V} = \frac{1}{6\pi^2} \left(\frac{\omega}{c} \right)^3 = \frac{1}{6\pi^2} (k_x^2 + k_y^2 + k_z^2)^{3/2} \quad - (10)$$

The total energy of the photons is:

$$U = \frac{E_0}{6\pi^2} \left(\frac{\omega}{c} \right)^3 L^3 \quad - (11)$$

If there is only one photon:

$$U = E_0 = \hbar \omega \quad - (12)$$

This result is obtained when the volume of radiation is:

$$V = L^3 = 6\pi^2 \left(\frac{c}{\omega} \right)^3 \quad - (13)$$

If the average energy of the black oscillator is denoted:

$$\langle \hbar \omega \rangle = \hbar \omega \frac{x}{1-x} \quad - (14)$$

$$U = \frac{\hbar \omega}{6\pi^2} \left(\frac{x}{1-x} \right) \left(\frac{\omega}{c} \right)^3 V \quad - (15)$$

and the energy density is :

$$\frac{U}{V} = \frac{\hbar \omega}{6\pi^2} \left(\frac{x}{1-x} \right) \left(\frac{\omega}{c} \right)^3 \quad - (16)$$

In the interval $d\omega$:

$$\frac{dU}{V} = \frac{\hbar \omega}{6\pi^2 c^3} \left(\frac{x}{1-x} \right) \left((\omega + d\omega)^3 - \omega^3 \right)$$

$$\sim \frac{\hbar \omega^3 d\omega}{2\pi^2 c^3} \left(\frac{x}{1-x} \right) \quad - (17)$$

In this interval :

$$\frac{dN}{V} = \frac{1}{6c^3 \pi^2} \left((\omega + d\omega)^3 - \omega^3 \right)$$

$$\sim \frac{1}{2} \frac{\omega^2}{c^3 \pi^2} d\omega \quad - (18)$$

It is usually considered that there are two states of polarization, so :

$$\frac{dN}{V} = \frac{\omega^2}{c^3 \pi^2} d\omega \quad - (19)$$

This is the usual Rayleigh Jeans density of states

4) This depends on the assumption of zero photon mass and no $\underline{b}^{(3)}$ field. The energy density in this case is:

$$\frac{u}{V} = \int \frac{\hbar \omega^3}{\pi^2 c^3} \left(\frac{x}{1-x} \right) d\omega \quad - (20)$$

and the intensity in watts per square metre is:

$$I = \frac{\hbar}{\pi^2 c^3} \int \omega^3 \left(\frac{x}{1-x} \right) d\omega \quad - (21)$$

where

$$x = \exp\left(-\frac{\hbar \omega}{kT} \right) \quad - (22)$$

The usual Stefan Boltzmann law is obtained.

where

$$I = \frac{\hbar}{\pi^2 c^3} \int_0^\infty \omega^3 \left(\frac{x}{1-x} \right) d\omega$$

$$= \left(\frac{\pi^2 \hbar^4}{60 c^2 \hbar^3} \right) T^4 \quad - (23)$$

The relation to the photon result:

$$E = \hbar \omega \quad - (24)$$

occurs where:

$$V = 6\pi^2 \left(\frac{1-x}{x} \right) \left(\frac{c}{\omega} \right)^3 \quad - (25)$$

where

$$u = \hbar \omega \quad - (26)$$

is eq. (15)