

### 326(3) : Rotational Quantization

Starting from the relativistic Lagrangian:

$$L = -\frac{mc^2}{\gamma} - U \quad (1)$$

Euler Lagrange equations give:

$$\frac{d}{dt}(\gamma m \dot{r}) - \gamma m r \dot{\theta}^2 = F(r) = -\frac{\partial U}{\partial r} \quad (2)$$

and

$$L = \gamma m r^2 \dot{\theta} = \gamma L_0 \quad (3)$$

is previous notation. The Lorentz factor is defined by

$$E^2 = \gamma^2 m^2 c^4 = p^2 c^2 + m^2 c^4 \\ = (H - U)^2 \quad (4)$$

and the potential energy is defined by:

$$U = H - \gamma mc^2 = H - \left(\frac{L}{L_0}\right)^2 mc^2 \quad (5)$$

Note that  $L$  is a constant of motion:

$$\frac{dL}{dt} = 0 \quad (6)$$

but  $L_0$  is not a constant of motion.

From eq. (4) :

$$H - mc^2 - U = \frac{c^2 p^2}{H + mc^2 - U} \quad (7)$$

2) In the non-relativistic approximation:

$$v \ll c \quad - (8)$$

and

$$\gamma \rightarrow 1. \quad - (9)$$

Furthermore, assume that:

$$H \sim E \quad - (10)$$

so

$$U \ll E. \quad - (11)$$

In these approximations:

$$H_1 = H - mc^2 \sim \frac{cp^2}{2mc^2 - U} + U \quad - (12)$$

$$= \frac{p^2}{2m} \left(1 - \frac{U}{2mc^2}\right)^{-1} + U$$

Now assume that:

$$U \ll 2mc^2 \quad - (13)$$

so:

$$H_1 = H - mc^2 \sim \frac{p^2}{2m} \left(1 + \frac{U}{2mc^2}\right) + U \quad - (14)$$

These approximations mean that:

$$E \sim mc^2 \text{ and } U \sim mc^2 \left(1 - \left(\frac{L}{L_0}\right)^2\right) \quad - (15)$$

$$L \rightarrow L_0 \quad - (16)$$

so, self consistently:

$$U \rightarrow 0 \quad - (17)$$

Eq. (14) quantizes to: - (18)

$$\left( -\frac{\nabla^2}{2m} + U \right) \psi = \left( H_1 + \frac{U}{4m^2 c^2} \nabla^2 \right) \psi$$

here

$$U = mc^2 \left( 1 - \left( \frac{L}{L_0} \right)^2 \right) \quad - (19)$$

$$\text{with } \left( \frac{L}{L_0} \right)^2 = \gamma^2 = \left( \frac{L}{mr^2 \dot{\theta}} \right)^2 \quad - (20)$$

However, it is known that:

$$U = -\frac{e^2}{4\pi \epsilon_0 r} \quad - (21)$$

so

$$1 - \left( \frac{L}{L_0} \right)^2 = -\frac{e^2}{4\pi \epsilon_0 mc^2 r} \quad - (22)$$

$$\text{and } \left( \frac{L}{L_0} \right)^2 = \left( \frac{L}{mr^2 \dot{\theta}} \right)^2 = 1 + \frac{e^2}{4\pi \epsilon_0 mc^2 r}$$

$$\rightarrow 1$$

$$- (23)$$

is the non-relativistic limit.

4) Eq. (23) gives:

$$\dot{\theta} = \frac{d\theta}{dt} = \frac{L}{mr^2} \left( 1 + \frac{e^2}{4\pi\epsilon_0 mc^2 r} \right)^{-1/2}$$
$$= \frac{L}{\gamma m r^2} \quad - (24)$$

Therefore

$$\dot{\theta} = \frac{d\theta}{dt} \sim \frac{L}{mr^2} \left( 1 - \frac{e^2}{8\pi\epsilon_0 mc^2 r} \right)^{-1/2} \quad - (25)$$
$$= \frac{L}{mr^2} \left( 1 - \frac{v^2}{c^2} \right)^{1/2}$$

if

$$v \ll c. \quad - (26)$$

Therefore

$$v^2 = \frac{e^2}{4\pi\epsilon_0 m r} \quad - (27)$$

This leads to an orbital equation using:

$$v^2 = \left( \frac{dr}{dt} \right)^2 + r^2 \left( \frac{d\theta}{dt} \right)^2 \quad - (28)$$

and:

$$\frac{d\theta}{dt} = \frac{d\theta}{dr} \frac{dr}{dt} \quad - (29)$$

From eq. (28):

$$\left(\frac{dr}{dt}\right)^2 + r^2 \left(\frac{d\theta}{dt}\right)^2 = \frac{e^2}{4\pi\epsilon_0 m r} \quad - (30)$$

$$\text{So: } \left(\frac{dr}{dt}\right)^2 = \frac{e^2}{4\pi\epsilon_0 m r} - r^2 \left(\frac{d\theta}{dt}\right)^2 \quad - (31)$$

- (32)

and

$$\frac{d\theta}{dr} = \frac{dt}{dr} \frac{d\theta}{dt} = \frac{d\theta}{dt} \left( \frac{e^2}{4\pi\epsilon_0 m r} - r^2 \left(\frac{d\theta}{dt}\right)^2 \right)^{-1/2}$$

i.e.

$$\frac{d\theta}{dr} = \frac{d\theta/dt}{\left( \frac{e^2}{4\pi\epsilon_0 m r} - r^2 \left(\frac{d\theta}{dt}\right)^2 \right)^{1/2}} \quad - (33)$$

where

$$\frac{d\theta}{dt} = \frac{L}{m r^2} \left( 1 - \frac{e^2}{8\pi\epsilon_0 m c^2 r} \right) \quad - (34)$$