

195(2): Derivation of the Orbital Equation from the Cotter Metric.

The starting Lagrangian is:

$$L = \frac{1}{2} mc^2 = \frac{m}{2} \left(c^2 A C^{1/2}(r) \left(\frac{dt}{d\tau} \right)^2 - B C^{1/2}(r) \left(\frac{dr}{d\tau} \right)^2 - C(r) \left(\frac{d\theta}{d\tau} \right)^2 \right) \quad (1)$$

The total energy is:

$$E = mc^2 A C^{1/2}(r) \frac{dt}{d\tau} \quad (2)$$

and the total angular momentum is:

$$L = m C(r) \frac{d\theta}{d\tau} \quad (3)$$

So:

$$E^2 = m^2 c^4 A^2 C(r) \left(\frac{dt}{d\tau} \right)^2 \quad (4)$$

$$L^2 = m^2 C^2(r) \left(\frac{d\theta}{d\tau} \right)^2 \quad (5)$$

Therefore:

$$mc^2 = \frac{1}{A C^{1/2}(r)} \frac{E^2}{mc^2} - m B C^{1/2}(r) \left(\frac{dr}{d\tau} \right)^2 - \frac{1}{m} \frac{L^2}{C(r)} \quad (6)$$

$$\frac{mc^2}{B C^{1/2}(r)} = \frac{1}{A B C(r)} \frac{E^2}{mc^2} - m \left(\frac{dr}{d\tau} \right)^2 - \frac{L^2}{m C^{3/2}(r) B} \quad (7)$$

$$m \left(\frac{dr}{d\tau} \right)^2 = \frac{1}{ABC(r)} \frac{E^2}{mc^2} - \frac{mc^2}{BC^{1/2}(r)} - \frac{L^2}{mBC^{3/2}(r)}$$

$$= \frac{1}{BC(r)} \left(\frac{E^2}{mc^2 A} - C^{1/2}(r) \left(mc^2 + \frac{1}{C(r)} \frac{L^2}{m} \right) \right) \quad - (8)$$

The equation of motion of Coulomb's metric is therefore:

$$\boxed{\frac{1}{2} m \left(\frac{dr}{d\tau} \right)^2 = \frac{1}{2BC(r)} \left(\frac{E^2}{mc^2 A} - C^{1/2}(r) \left(mc^2 + \frac{1}{C(r)} \frac{L^2}{m} \right) \right)} \quad - (9)$$

For spherical spacetime the same equation of motion is:

$$\frac{1}{2} m \left(\frac{dr}{d\tau} \right)^2 = \frac{1}{2} \left(\frac{E^2}{mc^2} - m(r) \left(mc^2 + \frac{L^2}{mr^2} \right) \right) \quad - (10)$$

To obtain the orbital equation of the Coulomb

metric we:

$$\frac{dr}{d\tau} = \frac{dr}{d\theta} \frac{d\theta}{d\tau}, \quad \frac{d\theta}{d\tau} = \frac{L}{mC(r)}, \quad - (11)$$

$$\text{So } \frac{dr}{d\tau} = \frac{L}{mC(r)} \frac{dr}{d\theta}, \quad m \left(\frac{dr}{d\tau} \right)^2 = \frac{L^2}{mC^2(r)} \left(\frac{dr}{d\theta} \right)^2 \quad - (12)$$

Therefore:

$$\frac{L^2}{C^2(r)} \left(\frac{dr}{d\theta} \right)^2 = \left(\frac{E^2}{mc^2 A} - C^{1/2}(r) \left(mc^2 + \frac{1}{C(r)} \frac{L^2}{m} \right) \right) \cdot \frac{1}{BC(r)} \quad - (13)$$

3) i.e

$$\left(\frac{dr}{d\theta}\right)^2 = \frac{mC(r)}{BL^2} \left(\frac{E^2}{mc^2 A} - C^{1/2}(r) \left(mc^2 + \frac{1}{C(r)} \frac{L^2}{m} \right) \right) \quad - (14)$$

The orbital equation is therefore:

$$\frac{dr}{d\theta} = \frac{m}{L} \left(\frac{C(r)}{B} \right)^{1/2} \left(\frac{E^2}{mc^2 A} - C^{1/2}(r) \left(mc^2 + \frac{1}{C(r)} \frac{L^2}{m} \right) \right)^{1/2} \quad - (15)$$

For a spherical spacetime the orbital equation is:

$$\frac{dr}{d\theta} = r^2 \left(\frac{1}{b^2} - m(r) \left(\frac{1}{a^2} + \frac{1}{r^2} \right) \right)^{1/2} \quad - (16)$$

$$\text{where } a = \frac{L}{mc}, \quad b = \frac{cL}{E} \quad - (17)$$

In order to calculate the deflection of light due to gravitation from the Cotton metric (15) the parameters E, L, B and $C(r)$ must be known. Therefore light deflection by gravitation cannot be calculated & as initial for any method based on an infinitesimal line element.