

196(4) : Calculation of the Precession Constant α

The precession constant α is calculated from:

$$\frac{1}{r} + \omega = \frac{d}{r} \left(\frac{x^2}{d} + \frac{1-x^2}{r} \right) \quad - (1)$$

of note 196(3), so:

$$\frac{d}{r} \left(\frac{rx^2 + d(1-x^2)}{dr} \right) = \frac{1 + \omega r}{r} \quad - (2)$$

and

$$rx^2 + d(1-x^2) = r(1 + \omega r) \quad - (3)$$
$$(r-d)x^2 = r(1 + \omega r) - d,$$
$$x^2 = \frac{r(1 + \omega r) - d}{r-d}$$

i.e.

$$x^2 = 1 + \frac{\omega r^2}{r-d} \quad - (4)$$

and the orbit is:

$$r = \frac{d}{1 + \epsilon \cos(x\theta)} \quad - (5)$$

In eq. (4), the quantities x , r and d can be determined by astronomy, so ω can be found experimentally.

The force law is:

$$F = -\frac{mMG}{r} \left(\frac{1}{r} + \omega \right) \quad - (6)$$

2) and in the classical limit:

$$\omega \rightarrow 0 \quad - (7)$$

so the spacetime becomes flat and:

$$F \rightarrow -\frac{2M_1 G}{r^2} \quad - (8)$$

In the limit (7), eq. (4) gives:

$$x^2 \rightarrow 1 \quad - (9)$$

and the orbit is a static ellipse:

$$r = \frac{d}{1 + e \cos \theta} \quad - (10)$$

The spin correction magnitude is given by:

$$\omega = -\frac{1}{r} + \frac{d}{r} \left(\frac{x^2}{d} + 1 - \frac{x^2}{r} \right) \quad - (11)$$

From eq. (11):

$$\omega \rightarrow 0 \quad \text{as } x \rightarrow 1 \quad - (12)$$

So tables of spin corrections can be drawn up for any object of mass m orbiting an object of mass M . This method can be used for any kind of orbit in cosmology.