

221(1) : Complete Analytical Solution of the Three-Particle Problem

As in previous notes the Lagrangian for the three particle problem is:

$$L = \frac{1}{2} (m_1 |\dot{\underline{r}}_1|^2 + m_2 |\dot{\underline{r}}_2|^2 + m_3 |\dot{\underline{r}}_3|^2) - \frac{m_1 m_2 G}{R_1} - \frac{m_1 m_3 G}{R_2} - \frac{m_2 m_3 G}{R_3} \quad (1)$$

where

$$R_1 = |\underline{r}_1 - \underline{r}_2| \quad (2)$$

$$R_2 = |\underline{r}_1 - \underline{r}_3| \quad (3)$$

$$R_3 = |\underline{r}_2 - \underline{r}_3| \quad (4)$$

Now use:

$$m_1 \underline{r}_1 + m_2 \underline{r}_2 = 0 \quad (5)$$

and

$$\underline{R}_1 = \underline{r}_1 - \underline{r}_2 \quad (6)$$

Therefore:

$$m_1 |\dot{\underline{r}}_1|^2 + m_2 |\dot{\underline{r}}_2|^2 = \mu_1 |\dot{\underline{R}}_1|^2 \quad (7)$$

where

$$\mu_1 = \frac{m_1 m_2}{m_1 + m_2} \quad (8)$$

The Lagrangian (1) therefore becomes:

$$L = \frac{1}{2} \left(\mu_1 |\dot{\underline{R}}_1|^2 + m_3 |\dot{\underline{r}}_3|^2 \right) - \frac{m_1 m_2 G}{R_1} - \frac{m_1 m_3 G}{R_2} - \frac{m_2 m_3 G}{R_3} \quad - (9)$$

where:

$$|\dot{\underline{R}}_1|^2 = \dot{R}_1^2 + R_1^2 \dot{\theta}_1^2 \quad - (10)$$

Now consider the Euler Lagrange equation:

$$\frac{\partial L}{\partial R_1} = \frac{d}{dt} \frac{\partial L}{\partial \dot{R}_1} \quad - (11)$$

and

$$\frac{\partial L}{\partial \theta_1} = \frac{d}{dt} \frac{\partial L}{\partial \dot{\theta}_1} \quad - (12)$$

These can be written as:

$$\frac{d^2}{d\theta_1^2} \left(\frac{1}{R_1} \right) + \frac{1}{R_1} = - \frac{\mu_1 R_1^2}{L_1^2} F_1(R_1) \quad - (13)$$

where

$$L_1 = \mu_1 R_1^2 \frac{d\theta_1}{dt} \quad - (14)$$

$$F_1 = - \frac{\partial u_1}{\partial R_1} \quad - (15)$$

and

$$U_1 = - \frac{m_1 m_2 G}{R_1} \quad - (16)$$

The solution of eq. (13) is:

$$R_1 = \frac{d_1}{1 + \epsilon_1 \cos \theta_1} \quad - (17)$$

where

$$L_1^2 = d_1 \mu_1 m_1 m_2 G \quad - (18)$$

and

$$\epsilon_1 = \left(1 + \frac{2E_1 L_1^2}{\mu_1 m_1^2 m_2^2 G^2} \right)^{1/2} \quad - (19)$$

Eq. (17) is an ellipse with major and minor semi-axes:

$$a_1 = \frac{d_1}{1 - \epsilon_1^2} = \frac{m_1 m_2 G}{2|E_1|} \quad - (20)$$

$$b_1 = \frac{d_1}{(1 - \epsilon_1^2)^{1/2}} = \frac{L_1}{(2\mu_1 |E_1|)^{1/2}} \quad - (21)$$

So:

$$\epsilon_1^2 = 1 - \left(\frac{b_1}{a_1} \right)^2 \quad - (22)$$

$$\begin{aligned} d_1 &= a_1 (1 - \epsilon_1^2) \quad - (23) \\ &= \frac{b_1^2}{a_1} \end{aligned}$$

4) Perseps d_1 and e_1 can be measured in astronomy. The actually measured quantities are a_1 and b_1 . They are related to the distance of closest approach, r_{min} , and furthest distance away, r_{max} , by:

$$r_{min} = a(1-e) \quad - (24)$$

$$r_{max} = a(1+e) \quad - (25)$$

Similarly:

$$R_2 = \frac{d_2}{1+e_2 \cos \theta_2}, \quad - (26)$$

$$L_2 = \mu_2 R_2^2 \frac{d\theta_2}{dt}, \quad - (27)$$

$$\mu_2 = \frac{m_1 m_3}{m_1 + m_3} \quad - (28)$$

$$e_2^2 = 1 - \left(\frac{b_2}{a_2}\right)^2, \quad d_2 = \frac{b_2^2}{a_2} \quad - (29)$$

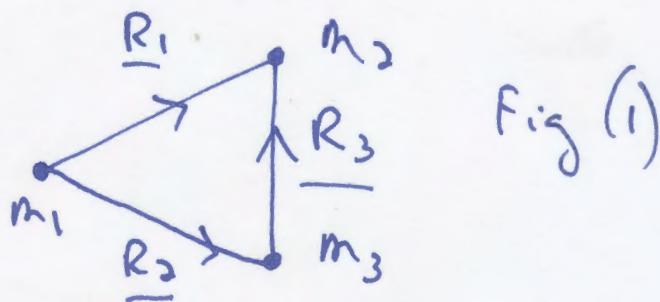
Thirdly:

$$R_3 = \frac{d_3}{1+e_3 \cos \theta_3}, \quad - (30)$$

$$L_3 = \mu_3 R_3^2 \frac{d\theta_3}{dt} \quad - (31)$$

$$\mu_3 = \frac{m_2 m_3}{m_2 + m_3}, \quad - (32)$$

$$\epsilon_3^2 = 1 - \left(\frac{b_3}{a_3}\right)^2, \quad \alpha_3 = \frac{b_3^2}{a_3} \quad - (33)$$



The three ellipses are related by :

$$\underline{R_1} = \underline{R_2} + \underline{R_3} \quad - (34)$$

The elapsed time for each ellipse are:

$$t_i = (1 - \epsilon_i^2)^{3/2} \left(\frac{\tau_i}{2\pi} \right) \int \frac{d\theta}{(1 + \epsilon_i \cos \theta)^3} \quad - (35)$$

where τ_i are the time for a particle on each ellipse to sweep out 2π radians.

In cylindrical polar coordinates:

$$\underline{R_1} = R_1 \underline{e_r} \quad - (36)$$

6) Similarly:

$$\underline{R}_2 = R_2 \underline{e}_r \quad - (37)$$

$$\underline{R}_3 = R_3 \underline{e}_r \quad - (38)$$

From Stokes' Theorem:

$$\oint \underline{R} \cdot d\underline{r} = \int \underline{\nabla} \times \underline{R} \cdot \underline{n} dA \quad - (39)$$

where $\underline{S} := \underline{\nabla} \times \underline{R} \quad - (40)$

In cylindrical polar coordinates, if:

$$\underline{F} = F_r \underline{e}_r + F_\theta \underline{e}_\theta + F_z \underline{k} \quad - (41)$$

then: - (42)

$$\underline{\nabla} \times \underline{F} = \left(\frac{1}{r} \frac{\partial F_z}{\partial \theta} - \frac{\partial F_\theta}{\partial z} \right) \underline{e}_r + \left(\frac{\partial F_r}{\partial z} - \frac{\partial F_z}{\partial r} \right) \underline{e}_\theta + \frac{1}{r} \left(\frac{\partial(r F_\theta)}{\partial r} - \frac{\partial F_r}{\partial \theta} \right) \underline{k}.$$

$$\text{So: } \underline{\nabla} \times \underline{R}_i = -\frac{1}{R_i} \frac{\partial R_i}{\partial \theta_i} \underline{k} \quad - (43)$$

$$i = 1, 2, 3$$

Therefore:

$$\oint \underline{R}_i \cdot d\underline{r} = - \int \frac{1}{R_i} \frac{\partial R_i}{\partial \theta_i} dA_i \quad - (44)$$

$i = 1, 2, 3$

In each case:

$$dA_i = \frac{1}{2} R_i^2 d\theta_i \quad - (45)$$

$i = 1, 2, 3$

Therefore:

$$\oint \underline{R_i} \cdot \underline{dr} = -\frac{1}{2} \int R_i dR_i \quad - (46)$$

and it follows that:

$$\boxed{\int R_1 dR_1 = \int R_2 dR_2 + \int R_3 dR_3} \quad - (47)$$

This equation relates the three ellipses.

For each ellipse:

$$R = \frac{d}{1 + e \cos \theta}, \quad \frac{dR}{d\theta} = \frac{e}{d} R^2 \sin \theta,$$
$$dR = \frac{e}{d} R^2 \sin \theta d\theta \quad - (48)$$

So:

$$\frac{e_1}{d_1} \int R_1^3 \sin \theta d\theta = \frac{e_2}{d_2} \int R_2^3 \sin \theta d\theta + \frac{e_3}{d_3} \int R_3^3 \sin \theta d\theta$$

$- (49)$

8)

i.e.

$$\boxed{\frac{\epsilon_1}{d_1} R_1^3 = \frac{\epsilon_2}{d_2} R_2^3 + \frac{\epsilon_3}{d_3} R_3^3} \quad - (50)$$

These results follow from the fact that:

$$\int R_1^3 \sin \theta_1 d\theta_1 = \int R_1^3 \sin \theta d\theta \quad - (51)$$

$$\int R_2^3 \sin \theta_2 d\theta_2 = \int R_2^3 \sin \theta d\theta \quad - (52)$$

$$\int R_3^3 \sin \theta_3 d\theta_3 = \int R_3^3 \sin \theta d\theta \quad - (53)$$
