

265(1) : Derivation of the Gravitational Red Shift.

Consider the metric of special relativity:

$$c^2 d\tau^2 = c^2 dt^2 - r^2 dt^2 - (1)$$

where

$$\sqrt{2} = \left(\frac{dr}{dt}\right)^2 + r^2 \left(\frac{d\theta}{dt}\right)^2, - (2)$$

so

$$\left(\frac{dr}{dt}\right)^2 = 1 - \frac{v^2}{c^2} - (3)$$

and

$$\frac{dt}{d\tau} = \gamma = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} - (4)$$

Eq. (4) defines the Lorentz factor in terms of

the velocity defined in eq. (2).
 The gravitational red shift can be defined by eq. (4) in terms of this velocity v . Note carefully that there is no need of the failed Einstein equation or Schwarzschild metric. It is a common misconception that the gravitational red shift is due to Einstein. In fact it can be derived straightforwardly from Newtonian dynamics as follows:

Consider the equivalence principle in the conventional form:

2)

$$\underline{F} = m \frac{d\underline{v}}{dt} = m \underline{g} = -m \frac{Mg}{r^2} \quad (5)$$

The work done on a particle by a force \underline{F} in transforming the particle from condition 1 to condition 2 is:

$$W_{1,2} = \int_1^2 \underline{F} \cdot d\underline{r} \quad (6)$$

in which:

$$\begin{aligned} \underline{F} \cdot d\underline{r} &= m \frac{d\underline{v}}{dt} \cdot \frac{d\underline{r}}{dt} dt = m \frac{d\underline{v}}{dt} \cdot \underline{v} dt \\ &= \frac{m}{2} \frac{d}{dt} (\underline{v} \cdot \underline{v}) dt = \frac{m}{2} \frac{d}{dt} (\underline{v}^2) dt \\ &= d\left(\frac{1}{2} m \underline{v}^2\right) \end{aligned} \quad (7)$$

So,

$$\int \underline{F} \cdot d\underline{r} = \frac{1}{2} m \underline{v}^2 \quad (8)$$

(Maria and Thornton, 3rd edition, page 71).

From eqns. (5) and (8):

$$\begin{aligned} \int \underline{F} \cdot d\underline{r} &= \frac{1}{2} m \underline{v}^2 = - \int \frac{m Mg}{r^2} dr \\ &= \underline{m Mg} \end{aligned} \quad (9)$$

So:

$$\boxed{\underline{v}^2 = \frac{2Mg}{r}} \quad (10)$$

3) The velocity referred in Eq. (10) is

$$v^2 = \left(\frac{dr}{dt} \right)^2 - (11)$$

The gravitational red shift can therefore be defined as:

$$\frac{dt}{d\tau} = \gamma = \left(1 - \frac{2Mg}{c^2 r} \right)^{-1/2} - (12)$$

In eq. (12), the old Schwarzschild radius is defined as:

$$r_0 = \frac{2Mg}{c^2} - (13)$$

and is a consequence of Newtonian dynamics, not of Einsteinian dynamics.

If $r_0 \ll r$ - (14)

then

$$\boxed{\frac{dt}{d\tau} \sim \frac{Mg}{c^2 r}} - (15)$$

which is the gravitational red shift, Q.E.D.

Note carefully that eq. (15) is an approximation, and is valid only if the velocity is defined as radial; eq. (11). This is the case when an object falls directly to the

4) Earth's surface. This was tested by the Harvard tower experiment, but it is a test of Newtonian dynamics incorporated into the Minkowski metric. It is not a test of the Einstein theory at all.

More generally \sqrt{v} is defined by eq. (2).

For example if: $r = \frac{d}{1 + \epsilon \cos \theta} \quad - (16)$

$$\sqrt{v^2} = M\bar{b} \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (17)$$

then:

where a is the semi-major axis of the ellipse:

$$a = \frac{d}{1 - \epsilon^2}. \quad - (18)$$

The gravitational red shift is then:

$$\frac{dt}{dr} = \gamma = \left(1 - \frac{M\bar{b}}{c^2} \left(\frac{2}{r} - \frac{1}{a} \right) \right)^{-1/2} \quad - (19)$$

The precessing ellipse observed in planetary orbits is:

$$r = \frac{d}{1 + \epsilon \cos(x\theta)} \quad - (20)$$

$$\text{where } x = 1 + \frac{3M\bar{b}}{c^2 d}. \quad - (21)$$

From eqs. (2) and (20):

$$v^2 = \alpha^2 M G \left(\frac{2}{r} - \frac{1}{a} \right) + \frac{L^2}{m^2 r^2} (1 - \alpha^2) \quad -(22)$$

where

$$L^2 = \alpha m^2 M G \quad -(23)$$

so

$$\begin{aligned} v^2 &= \alpha^2 M G \left(\frac{2}{r} - \frac{1}{a} \right) + \alpha \frac{M G}{r^2} (1 - \alpha^2) \\ &= M G \left[\alpha^2 \left(\frac{2}{r} - \frac{1}{a} \right) + \frac{\alpha}{r^2} (1 - \alpha^2) \right] \end{aligned} \quad -(24)$$

The expected gravitational red shift in the experimentally observed precessing orbit is :

$$\frac{t}{\tau} = \gamma = \left(1 - \frac{M G}{c^2} \left[\alpha^2 \left(\frac{2}{r} - \frac{1}{a} \right) + \frac{\alpha}{r^2} (1 - \alpha^2) \right] \right)^{-1/2} \quad -(25)$$

where

$$\alpha = 1 + \frac{3 M G}{c^2 \alpha} \quad -(26)$$

This result can be tested experimentally.