

238(5) : Hamiltonian Method of Calculating $(p/L)^2$
and the Special Relativistic Orbit.

Consider the Hamiltonian method based on:

$$H = \gamma mc^2 + U \quad - (1)$$

of special relativity. It follows that:

$$H = (p^2 c^2 + m^2 c^4)^{1/2} + U \quad - (2)$$

$$\begin{aligned} \text{so: } H_0 = H - mc^2 &= \frac{p^2 c^2}{H - U + mc^2} + U \quad - (3) \\ &= \frac{p^2 c^2}{E + mc^2} + U \end{aligned}$$

Here:

$$E = \gamma mc^2, \quad p = \gamma m v_0 \quad - (4)$$

where

$$v_0^2 = \left(\frac{dr}{dt} \right)^2 + r^2 \left(\frac{d\theta}{dt} \right)^2 \quad - (5)$$

So:

$$H_0 = \left(\frac{\gamma^2}{1 + \gamma} \right) m v_0^2 + U \quad - (6)$$

$$= \left(\frac{E^2}{m c^2 (E + m c^2)} \right) m v_0^2 + U$$

$$\xrightarrow{v_0 \ll c} \quad \frac{1}{2} m v_0^2 + U$$

2) In eq. (5):

$$L_0 = m r^2 \frac{d\theta}{dt} \quad - (7)$$

is the non relativistic angular momentum, a constant of motion. Also:

$$\frac{dr}{dt} = \frac{dr}{d\theta} \frac{d\theta}{dt} \quad - (8)$$

It follows that:

$$H_0 - U = \left(\frac{E^2}{m c^2 (E + m c^2)} \right) \frac{L_0^2}{m r^4} \left(\left(\frac{dr}{d\theta} \right)^2 + r^2 \right) \quad - (9)$$

and

$$\left(\frac{dr}{d\theta} \right)^2 + r^2 = m r^4 \left(\frac{H_0 - U}{L_0^2} \right) \left(\frac{m c^2 (E + m c^2)}{E^2} \right)$$

$$\begin{aligned} &\xrightarrow{v \ll c} 2 m r^4 \left(\frac{H_0 - U}{L_0^2} \right) \quad - (10) \\ &= m r^4 \left(\frac{p_0}{L_0} \right)^2 \end{aligned}$$

This is the same as the equation derived from

the metric

$$c^2 d\tau^2 = (c^2 - v^2) dt^2 \quad - (11)$$

therefore the transition from classical to relativistic theory is given by:

$$3) \quad T = \frac{1}{2} \frac{p_0^2}{m} \rightarrow \left(\frac{E^2}{mc^2(E+mc^2)} \right) \frac{p_0^2}{m} \quad - (11)$$

i.e.

$$p_0^2 \rightarrow 2 \left(\frac{E^2}{mc^2(E+mc^2)} \right) p_0^2 \quad - (12)$$

$$\text{So } \left(\frac{p_0}{L_0} \right)^2 \rightarrow 2 \left(\frac{E^2}{mc^2(E+mc^2)} \right) \left(\frac{p_0}{L_0} \right)^2 \quad - (13)$$

The relativistic orbit is therefore given by :

$$\left(\frac{dr}{dt} \right)^2 + r^2 = 2r^4 \left(\frac{E^2}{mc^2(E+mc^2)} \right) \left(\frac{p_0}{L_0} \right)^2 \quad - (14)$$

In this equation :

$$p_0^2 = m^2 MG \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (15)$$

and

$$L_0^2 = m^2 \underline{M} G d \quad - (16)$$

$$\text{So } \left(\frac{p_0}{L_0} \right)^2 = \frac{1}{d} \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (17)$$

where d is the half right latitude and a the semi major axis. These two observables are

3) related by: $a = \frac{d}{1-\epsilon^2} - (18)$

The classical orbit is:

$$\begin{aligned} \left(\frac{dr}{d\theta}\right)^2 &= r^4 \left(\frac{p_0}{L_0}\right)^2 - r^2 \\ &= \frac{r^4}{d} \left(\frac{2}{r} - \frac{1}{a}\right) - r^2 \quad - (19) \\ &= \frac{2r^3}{d} - \frac{r^4}{d^2} (1-\epsilon^2) - r^2 \end{aligned}$$

It can be checked that eq. (19) is the same as the given by: $r = \frac{d}{1+\epsilon \cos \theta}$, $\cos \theta = \frac{1}{\epsilon} \left(\frac{d}{r} - 1\right) - (20)$

for which:

$$\begin{aligned} \left(\frac{dr}{d\theta}\right)^2 &= \frac{\epsilon^2 r^4}{d^2} \sin^2 \theta = \frac{\epsilon^2 r^4}{d^2} (1 - \cos^2 \theta) - (21) \\ &= \frac{\epsilon^2 r^4}{d^2} \left(1 - \frac{1}{\epsilon^2} \left(\frac{d}{r} - 1\right)^2\right) \\ &= \frac{\epsilon^2 r^4}{d^2} - r^2 + \frac{2r^3}{d} - \frac{r^4}{d^2} \end{aligned}$$

Eq. (19) and (21) are the same, Q.E.D.

4) So the relativistic orbit is given by:

$$\left(\frac{dr}{dt}\right)^2 + r^2 = \frac{2r^4}{d} \left(\frac{E^2}{mc^2(E+mc^2)} \right) \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (22)$$

where $E = \gamma mc^2 = \left(1 - \frac{v_0^2}{c^2}\right)^{-1/2} mc^2 \quad - (23)$

and $v_0^2 = \frac{mG}{d} \left(\frac{2}{r} - \frac{1}{a} \right) \quad - (24)$

Eg. (22) is the exact orbit of special relativity, GED.

From numerical work it is known that it is a precessing ellipse (UFT 324 and UFT 325).
It can be compared with the theory:

$$r = \frac{d}{1 + \epsilon \cos(x\theta)} \quad - (25)$$

and $\frac{dr}{dt} = \frac{x \epsilon r^2 \sin(x\theta)}{d} \quad - (26)$

in order to find x of special relativity,
and also the general precessing ellipse:

$$r = \frac{d}{1 + \epsilon \cos(\theta_1(\theta))} \quad - (27)$$